

Magnetic behavior of a nonextensive S -spin system: Possible connections to manganites

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We analyzed the magnetic behavior of a S -spin system within the framework of Tsallis nonextensive statistics, employing the normalized approach. Unusual properties on magnetization, entropy, and susceptibility emerge as a consequence of the nonextensivity. We further show that the nonextensive approach can be relevant to the field of manganites, materials which exhibit long-range interactions and fractality, two basic ingredients for nonextensivity. Our results are in qualitative agreement with experimental data in $\text{La}_{0.67}\text{Ca}_{0.33}\text{MnO}_3$ and $\text{Pr}_{0.5}\text{Ca}_{0.5}\text{Mn}_{0.95}\text{Ga}_{0.05}\text{O}_3$ manganites.

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I. INTRODUCTION

In 1988, Tsallis¹ proposed a q -dependent entropy functional which generalized the standard Maxwell-Boltzmann definition, to include nonextensive systems. In this formalism, “ q ” is called the *entropic index*, and measures the degree of nonextensivity of the system. A few years later, Curado and Tsallis² revised the formalism introducing the *unnormalized* constraint to the internal energy, and showed that from this entropy functional a nonextensive thermodynamics could be derived. This generalization included the classical thermodynamics (Maxwell-Boltzmann), which could be recovered when q is set to unity.

After the work of Curado and Tsallis, this formalism has been successfully applied to various physical systems, where the Maxwell-Boltzmann framework fails. These include self-gravitating systems,^{3,4} turbulence,⁵⁻⁸ anomalous diffusion,⁹⁻¹² velocities of galaxies,¹³ solar neutrinos,¹⁴ etc.

In spite of these successes, some drawbacks were identified in the early formalism, namely, (a) the density operator was not invariant under a uniform translation of energy spectrum; (b) the q -expected value of the identity operator was not the unity; and (c) energy was not conserved.¹⁵ In a more recent work, Tsallis *et al.*¹⁵ circumvented these difficulties by introducing the *normalized* constraint to the internal energy of the system.

Concerning applications to magnetic systems, the unnormalized formalism was used by Portesi *et al.*¹⁶ and Nobre and Tsallis,¹⁷ to describe the paramagnetic behavior of a system with N spins $1/2$. The authors found a non-measurable magnetic susceptibility, since it was exponentially dependent on the number N of particles. Martinez *et al.*¹⁸ analyzed the same $S=1/2$ system in the framework of the normalized formalism,¹⁵ finding a similar result to that reported by the Portesi *et al.* and Nobre and Tsallis.

In the present work, we analyzed the paramagnetic behavior of N spins S , within the normalized formalism, and showed that the system effective temperature T does not relate with the inverse of the Lagrange multiplier β , as it is normally assumed, but with the inverse of a re-scaled parameter β^* ($T=1/k\beta^*$), already introduced by Tsallis *et al.*¹⁵ In

this approach, the nonphysical result found in Refs. 16,17, and 18 was no longer found, and the susceptibility becomes proportional to the number of particles, as experimentally expected. In those early works, no attempt was made to correlate theoretical results to experimental magnetic systems of any kind.

We also suggest that the manganites are physical systems where the present concepts can be tested. Experimental data of Amaral *et al.*¹⁹ and Hebert *et al.*²⁰ are in qualitative agreement with the results here reported. An analysis of the nonextensivity in manganites in the ferromagnetic phase, also taking $T=1/k\beta^*$ as an effective temperature, was already published elsewhere.²¹

II. MODEL

We consider the Hamiltonian of a single spin S in a static and homogeneous magnetic field:

$$\hat{\mathcal{H}} = -\hat{\boldsymbol{\mu}} \cdot \vec{B} = -\hat{\mu}_z B = -g\mu_B \hat{S}_z B, \quad (1)$$

where $g=2$ for $J=S$. The q -generalized magnetic moment thermal average $\langle \hat{\mu}_z \rangle_q$ is

$$\langle \hat{\mu}_z \rangle_q = g\mu_B \langle \hat{S}_z \rangle_q = g\mu_B S B_S^{(q)}, \quad (2)$$

where $\langle \hat{S}_z \rangle_q$ is q -generalized thermal average spin operator and $B_S^{(q)}$ is the *generalized Brillouin function*. This first quantity can be determined in the Tsallis framework of normalized q -expectation values,¹⁵ from

$$\langle \hat{S}_z \rangle_q = \frac{\text{Tr}\{\hat{\rho}^q \hat{S}_z\}}{\text{Tr}\{\hat{\rho}^q\}}, \quad (3)$$

where $\hat{\rho}$ is the density operator, derived from the maximization of the entropy. Below, we analyze two different proposals for the entropy and show that rescaling the Lagrange parameter, the same density operator emerges from both definitions.

A. Tsallis entropy

The entropy of the system is defined as^{1,2,15}

$$\mathcal{S}_q = \text{Tr}\{\hat{\rho}^q \hat{\mathcal{S}}_q\} = \frac{k}{(q-1)} [1 - \text{Tr}\{\hat{\rho}^q\}], \quad (4)$$

where

$$\hat{\mathcal{S}}_q = -k \ln_q \hat{\rho}, \quad (5)$$

and k a positive constant. The generalized logarithm is defined as²²

$$\ln_q f = (1-q)^{-1} (f^{1-q} - 1). \quad (6)$$

The corresponding density operator $\hat{\rho}$ can be determined from the maximization of the \mathcal{S}_q functional, subjected to a q -normalized constraint,¹⁵

$$U_q = \frac{\text{Tr}\{\hat{\rho}^q \hat{\mathcal{H}}\}}{\text{Tr}\{\hat{\rho}^q\}}, \quad (7)$$

and the normalization of density operator $\text{Tr}\{\hat{\rho}\} = 1$, leading to

$$\hat{\rho} = \frac{1}{Z_q} \left[1 - (1-q) \frac{\beta}{\text{Tr}\{\hat{\rho}^q\}} (\hat{\mathcal{H}} - U_q) \right]^{1/(1-q)}, \quad (8)$$

where Z_q is the generalized partition function,

$$Z_q = \text{Tr} \left\{ \left[1 - (1-q) \frac{\beta}{\text{Tr}\{\hat{\rho}^q\}} (\hat{\mathcal{H}} - U_q) \right]^{1/(1-q)} \right\}, \quad (9)$$

and β is the Lagrange parameter associated with internal energy constraint.^{15,23} In this case, the entropy has a well defined concavity, for any value of q , being concave for $q > 0$, and convex for $q < 0$.^{1,2,15,23}

B. Normalized Tsallis entropy

If the entropy functional $\hat{\mathcal{S}}_q$ is defined as proposed by Rajagopal and Abe,²⁴

$$\mathcal{S}_q = \frac{\text{Tr}\{\hat{\rho}^q \hat{\mathcal{S}}_q\}}{\text{Tr}\{\hat{\rho}^q\}} = \frac{k}{(q-1)} \left[\frac{1}{\text{Tr}\{\hat{\rho}^q\}} - 1 \right], \quad (10)$$

the density operator $\hat{\rho}$ can be determined in a similar way as above, yielding

$$\hat{\rho} = \frac{1}{Z_q} [1 - (1-q) \beta \text{Tr}\{\hat{\rho}^q\} (\hat{\mathcal{H}} - U_q)]^{1/(1-q)}, \quad (11)$$

$$Z_q = \text{Tr}\{[1 - (1-q) \beta \text{Tr}\{\hat{\rho}^q\} (\hat{\mathcal{H}} - U_q)]^{1/(1-q)}\}. \quad (12)$$

In this case, the q parameter is restricted to the interval $0 \leq q \leq 1$, preserving the entropy concavity.²⁴

The density operators emerging from both scenarios [Eqs. (8) and (11)] can be rewritten in the forms

$$\hat{\rho} = \frac{1}{Z'_q} [1 - (1-q) \beta^* \hat{\mathcal{H}}]^{1/(1-q)}, \quad (13)$$

$$Z'_q = \text{Tr}\{[1 - (1-q) \beta^* \hat{\mathcal{H}}]^{1/(1-q)}\}, \quad (14)$$

where

$$\beta^* = \frac{\beta}{\text{Tr}\{\hat{\rho}^q\} + (1-q) U_q \beta}, \quad (15)$$

for case A, and

$$\beta^* = \frac{\beta}{\frac{1}{\text{Tr}\{\hat{\rho}^q\}} + (1-q) U_q \beta} \quad (16)$$

for case B. Here we suggest that the effective temperature is

$$T = \frac{1}{k \beta^*}, \quad (17)$$

and the density operator becomes independent of the initial entropy functional.

The magnetic behavior of a S -spin system will be analyzed as a function of the parameter $x^* = g \mu_B S B \beta^*$. In fact, in terms of x^* , the generalized Brillouin function is given by

$$B_S^{(q)} = \frac{1}{S} \langle \hat{\mathcal{S}}_z \rangle_q = \frac{1}{S} \frac{\sum_{m_s=-S}^{+S} m_s \left[1 + (1-q) x^* \frac{m_s}{S} \right]^{q/(1-q)}}{\sum_{m_s=-S}^{+S} \left[1 + (1-q) x^* \frac{m_s}{S} \right]^{q/(1-q)}}. \quad (18)$$

It is to be remarked that cutoff procedure^{23,25-27} implies that those states that do not satisfy the condition

$$1 + (1-q) x^* \frac{m_s}{S} \geq 0 \quad (19)$$

must be excluded from the summation. In other words, these states are assigned a zero probability amplitude, preserving the positive definition of the density operator.¹⁶

For a system constituted by N spins- S particles, the q -generalized spin operator thermal average [Eq.(3)] can be written as

$$\langle \hat{\mathcal{S}}_z \rangle_{q,N} = \frac{\sum_{m_s=-NS}^{+NS} Y(m_s) m_s \left[1 + (1-q) x^* \frac{m_s}{S} \right]^{q/(1-q)}}{\sum_{m_s=-NS}^{+NS} Y(m_s) \left[1 + (1-q) x^* \frac{m_s}{S} \right]^{q/(1-q)}}, \quad (20)$$

where $Y(m_s)$ is the multiplicity and

$$\sum_{m_s=-NS}^{NS} Y(m_s) = (2S+1)^N. \quad (21)$$

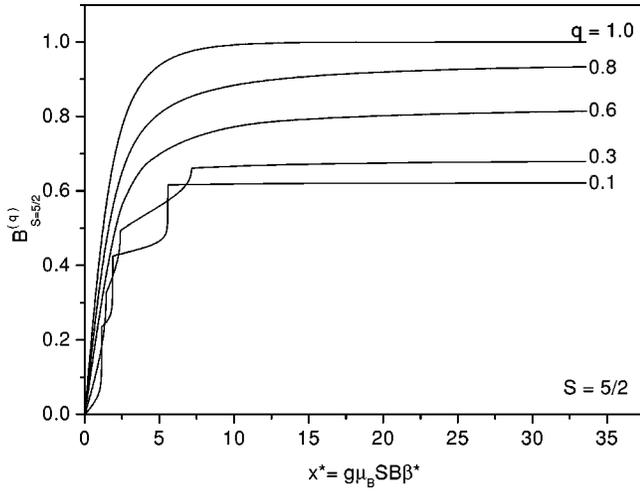


FIG. 1. Generalized Brillouin function for several q values and $S=5/2$, as a function of x^* .

In the particular case of N spin-1/2 particles, $Y(m_s)$ is simply given by

$$Y(m_s) = \frac{N!}{\left(\frac{N}{2} - m_s\right)! \left(\frac{N}{2} + m_s\right)!}. \quad (22)$$

III. RESULTS AND DISCUSSION

When q is different from unity, general expressions for magnetic observables are difficult to calculate and interpret, even for simple systems. This hinders the comparison between theoretical predictions and experimental results. Numerical methods, on the other hand, provide a means to directly calculate observables, allowing a deeper understanding of the theoretical results.

Figure 1 displays the generalized Brillouin function $B_S^{(q)}$ vs x^* [Eq. (18)], for different values of q and $S=5/2$. For q up to 0.5, a series of kinks appear in the curve. One also notes that the saturation value $B_S^{(q)}|_{\text{SAT}}$ decreases with decreasing q . This is more clearly shown in Figs. 2(a) and 2(b), where $B_S^{(q)}|_{\text{SAT}}$ is plotted as a function of q , for half-integer and integer spin values, respectively. The behavior of a classical spin is also included and can be exactly calculated from Eq. (18) with $S \rightarrow \infty$, giving

$$B_S^{(q)}|_{\text{SAT}} = \frac{1}{2-q}. \quad (23)$$

The occupation probability (OP), as a function of x^* , for each energy level of $S=5/2$, are displayed in Figs. 3(a)–3(c), for several values of q . Figure 3(d) shows the same quantity for $S=2$ and $q=0.1$. We observe that the OP does not vanish for negative energy levels ($m_s > 0$), even for very large values of x^* , in contrast to what happens in the case $q=1$ (Maxwell-Boltzmann). From Fig. 3(c) we can see that, for $q=0.1$, the OP of the positive energy states ($m_s < 0$) vanish sharply at the same x^* values as the kinks observed in Fig. 1. This occurs for $q < 0.5$. Therefore, we correlate the kinks

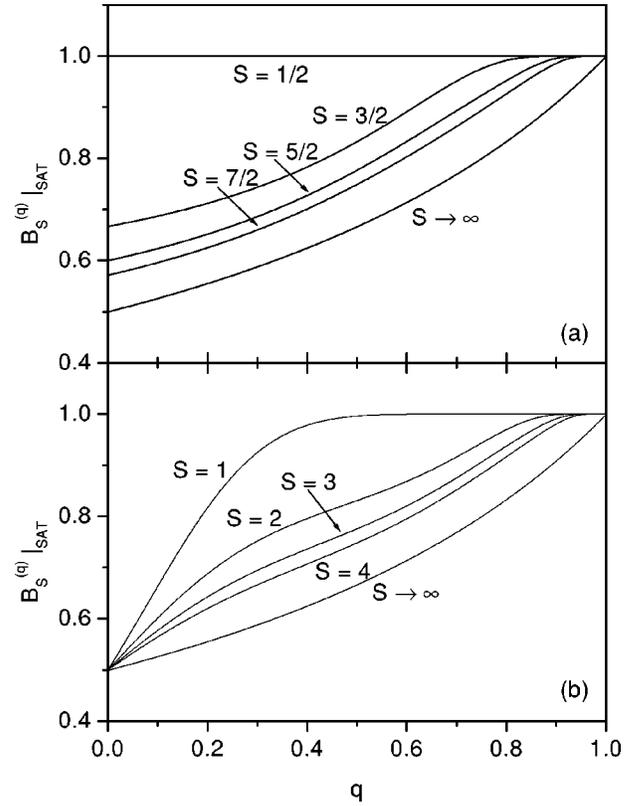


FIG. 2. Saturation magnetization $B_S^{(q)}|_{\text{SAT}}$ as a q function, for (a) half-integer and (b) integer S values.

observed on the magnetization curves to the loss of occupation of the most energetic states. This is a consequence of the Tsallis cutoff. In fact, from Eq. (19), the x^* values where the kinks occurs can be derived, following

$$x_{m_s}^{\text{kinks}} = \frac{S}{|m_s|(1-q)}. \quad (24)$$

It is to be remarked that for a half-integer spin, $S+1/2$ kinks occur, and S for an integer spin.

Figures 4(a) and 4(b) display, respectively, the unnormalized [Eq. (4)] and normalized [Eq. (10)] entropy S_q , for $S=5/2$ and different q values. Note that, for $q < 0.5$, the kinks discussed before are present. For any $q < 1$, the entropy does not vanish in the limit $x^{*-1} \rightarrow 0$. In other words, even for a high field and/or a low temperature, the system has a finite entropy, which prevents a fully magnetized state. These features are valid for both entropy functionals (Sec. II).

A general expression for the magnetic susceptibility can be deduced from Eqs. (2) and (18),

$$\chi_q = \lim_{B \rightarrow 0} \left[\frac{\partial \langle \hat{\mu}_z \rangle_q}{\partial B} \right] = \frac{C^{(q)}}{T}, \quad (25)$$

where $C^{(q)} = qC^{(1)}$ is the generalized Curie constant. Figure 5 displays χ_q^{-1} as a function of x^{*-1} for $q=1.0$ and 0.8 , with $S=5/2$.

In order to compare the predictions of the proposed model to experimental data, we must investigate how the

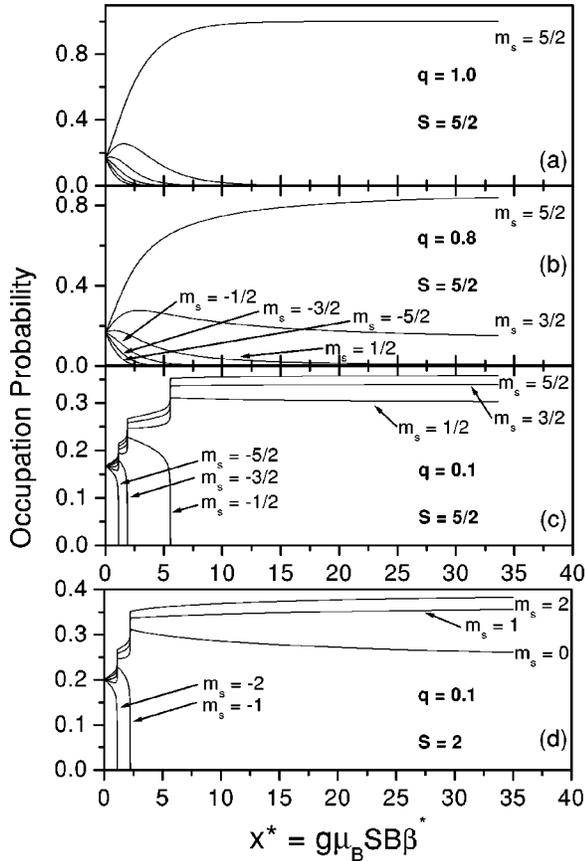


FIG. 3. Probability of the occupation of each energy level, for (a)–(c) $S=5/2$ and (d) $S=2$, as a function of x^* , and various values of q .

magnetic observables discussed before scales upon increasing the number of particles in the system. An extensive quantity \mathcal{F} scales as

$$\mathcal{F}_N = N\mathcal{F}_1, \quad (26)$$

where N is the system number of particle.

Particularly useful for the purpose of comparison, is the magnetic susceptibility. Portesi *et al.*¹⁶ and Nobre and Tsallis,¹⁷ considered the generalized magnetization on the paramagnetic phase of a N spin-1/2 system in the unnormalized formalism.^{2,15–17,23} They found, for χ_q ,

$$\chi_q = \frac{C_{S=1/2}^{(q)}}{T} 2^{N(1-q)}, \quad (27)$$

where

$$C_{S=1/2}^{(q)} = \frac{(g\mu_B)^2}{4k} Nq, \quad (28)$$

and $T=1/k\beta$ is the temperature of the system. Here β is the usual Lagrange parameter associated to the non-normalized constraint of internal energy.

Therefore, as N increases, the magnetic susceptibility becomes infinity for $q < 1$ and zero for $q > 1$. In other words, there is no linearity among χ_q and N . This phenomenon is

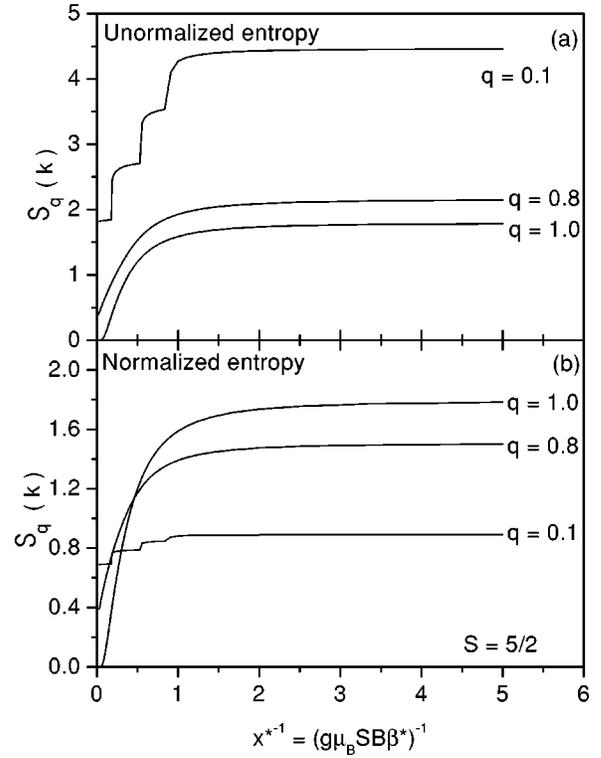


FIG. 4. (a) Unnormalized [Eq. (4)] and (b) normalized [Eq. (10)] entropy S_q , for $S=5/2$ and different q values, as a function of x^{*-1} .

called *dark magnetism*, in analogy to the cosmological concept of *dark matter* (see Ref. 16, and references therein). Martinez *et al.*¹⁸ analyzed the paramagnetic behavior of the same N spin-1/2 system within the normalized formalism,^{15,23} and found a result similar to that of Portesi and Nobre for χ_q .

Our proposal is that the paramagnetic behavior of a non-extensive N spin-1/2 system should be analyzed in the normalized formalism, using the density operator described in Eq. (13), taking β^* , instead of β , inversely proportional to the system temperature T . By doing so, the generalized paramagnetic susceptibility becomes

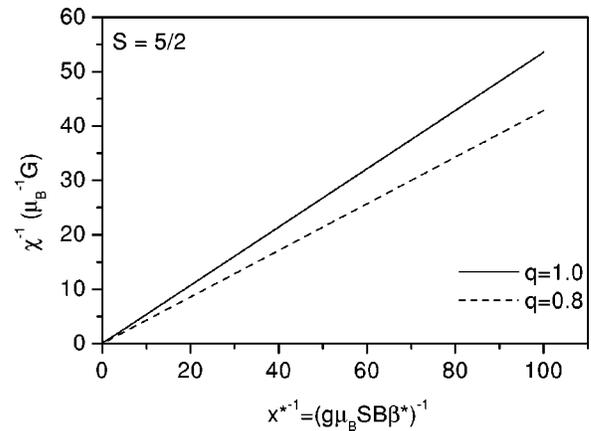


FIG. 5. Inverse of susceptibility for $q=1.0$ and 0.8 as a x^{*-1} function, for $S=5/2$.

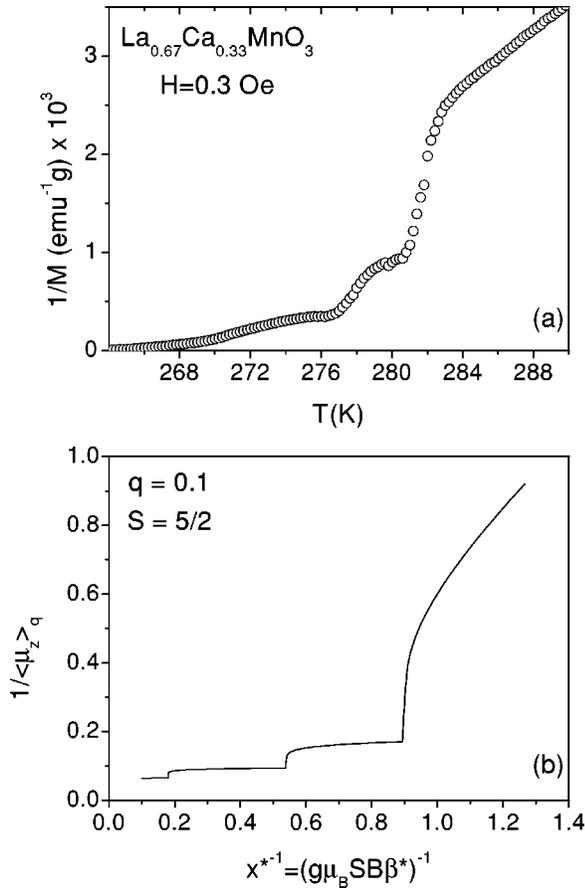


FIG. 6. (a) Temperature dependence of the inverse susceptibility for the manganite $\text{La}_{2/3}\text{Ca}_{1/3}\text{MnO}_3$, above the Curie temperature (267 K), at a low magnetic field. (b) Inverse of the magnetization [Eq. (18)] as a function of x^{*-1} , for $q=0.1$.

$$\chi_q = \frac{C_{S=1/2}^{(q)}}{T}, \quad (29)$$

which does not diverge as N increases, and is indeed proportional to N . In addition, this result is independent of the choice for the functional entropy.

IV. POSSIBLE CONNECTION TO EXPERIMENTAL RESULTS

Amaral *et al.*¹⁹ discussed the magnetic behavior of manganese oxides, namely, $\text{La}_{0.67}\text{Ca}_{0.33}\text{MnO}_3$ and verified steps on the curve of M^{-1} vs T in the paramagnetic phase, as shown in Fig. 6(a). These steps are analogs to those shown in Fig. 6(b), that also represents the inverse of magnetization as a function of x^{*-1} , for $q=0.1$, and encourage the idea of manganites as nonextensive objects.²¹ The authors took the change in the slope of the curve as an indication of cluster formation, which changes the effective moment of Mn ions. These clusters could give rise to fractal structures, as discussed by Dagotto *et al.*,²⁸ and therefore are in accordance to the ideas discussed here. In addition, Hebert *et al.*²⁰ found magnetization curves in $\text{Pr}_{0.5}\text{Ca}_{0.5}\text{Mn}_{0.95}\text{Ga}_{0.05}\text{O}_3$ that are also in qualitative agreement to those curves in Fig. 1.

V. CONCLUSION

In summary, we investigated the properties of a paramagnetic S -spin system under the Tsallis generalized statistics, on the normalized formalism. For $q < 0.5$, a series of kinks appears in the magnetization and entropy. This effect is a direct consequence of a peculiar occupation probability, as a result of the Tsallis cutoff, where the positive energy states ($m_s < 0$) vanish sharply. Additionally, the negative energy states ($m_s > 0$) share a nonzero occupation probability, preventing a fully magnetized state and the saturation magnetization decreases with decreasing q . We present evidences based on experimental results of Amaral *et al.*¹⁹ and Hebert *et al.*,²⁰ which add to and support our previous publication,²¹ where manganites were suggested to be magnetically nonextensive objects.

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¹C. Tsallis, *J. Stat. Phys.* **52**, 479 (1988).

²E. M. F. Curado and C. Tsallis, *J. Phys. A* **24**, L69 (1991).

³A. R. Plastino and A. Plastino, *Phys. Lett. A* **174**, 384 (1993).

⁴J. J. Aly, in *Proceedings of N-Body Problems and Gravitational Dynamics*, edited by F. Combes and E. Athanassoula (Publications de L'Observatoire de Paris, Paris, 1993), p. 19.

⁵T. Arimitsu and N. Arimitsu, *Phys. Rev. E* **61**, 3237 (2000).

⁶T. Arimitsu and N. Arimitsu, *J. Phys. A* **33**, L235 (2000).

⁷C. Beck, *Physica A* **277**, 115 (2000).

⁸C. Beck, G. S. Lewis, and H. L. Swinney, *Phys. Rev. E* **63**, 035303 (2001).

⁹A. R. Plastino and A. Plastino, *Physica A* **222**, 347 (1995).

¹⁰C. Tsallis and D. J. Bukman, *Phys. Rev. E* **54**, R2197 (1996).

¹¹M. Bologna, C. Tsallis, and P. Grigolini, *Phys. Rev. E* **62**, 2213 (2000).

¹²L. C. Malacarne, I. T. Pedron, R. S. Mendes, and E. K. Lenzi, *Phys. Rev. E* **63**, 030101(R) (2001).

¹³A. Lavagno, G. Kaniadakis, M. Rego-Monteiro, P. Quarati, and C. Tsallis, *Astrophys. Lett. Comm.* **35/6**, 449 (1998).

¹⁴A. Lavagno and P. Quarati, *Nucl. Phys. B* **87**, 209 (2000).

¹⁵C. Tsallis, R. S. Mendes, and A. R. Plastino, *Physica A* **261**, 534 (1998).

¹⁶M. Portesi, A. Plastino, and C. Tsallis, *Phys. Rev. E* **52**, R3317 (1995).

¹⁷F. R. Nobre and C. Tsallis, *Philos. Mag. B* **73**, 745 (1996).

¹⁸S. Martinez, F. Pennini, and A. Plastino, *Physica A* **282**, 193 (2000).

¹⁹V. S. Amaral, J. P. Araújo, Y. G. Pogorelov, J. M. B. Lopes dos Santos, P. B. Tavares, A. A. C. S. Lourenco, J. B. Sousa, and J. M. Vieira, *J. Magn. Magn. Mater.* **226**, 837 (2001).

- ²⁰S. Hebert, V. Hardy, A. Maignan, R. Mahendiran, M. Hervieu, C. Martin, and B. Raveau, *J. Solid State Chem.* **165**, 6 (2002).
- ²¹M. S. Reis, J. C. C. Freitas, M. T. D. Orlando, E. K. Lenzi, and I. S. Oliveira, *Europhys. Lett.* **58**, 42 (2002).
- ²²E. P. Borges, *J. Phys. A* **31**, 5281 (1998).
- ²³C. Tsallis, in *Nonextensive Statistical Mechanics and Its Applications*, edited by S. Abe and Y. Okamoto (Springer-Verlag, Heidelberg, 2001).
- ²⁴A. K. Rajagopal and S. Abe, *Phys. Rev. Lett.* **83**, 1711 (1999).
- ²⁵A. R. Lima and T. J. P. Penna, *Phys. Lett. A* **256**, 221 (1999).
- ²⁶A. Plastino and A. R. Plastino, *Braz. J. Phys.* **29**, 50 (1999).
- ²⁷C. Tsallis, *Braz. J. Phys.* **29**, 1 (1999).
- ²⁸E. Dagotto, T. Hotta, and A. Moreo, *Phys. Rep.* **344**, 1 (2001).